

Materials discovery for solar thermochemical water splitting applications using computational techniques

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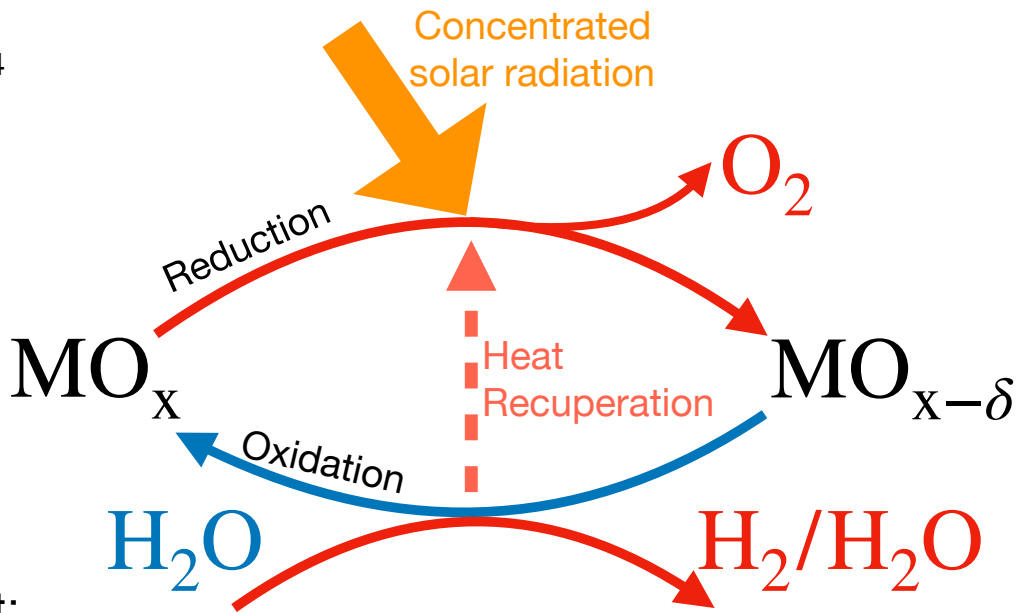
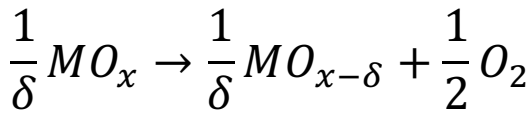
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Solar thermochemical (STC) production of H₂ and/or CO: oxide perovskites are potential candidates

Candidates so far:

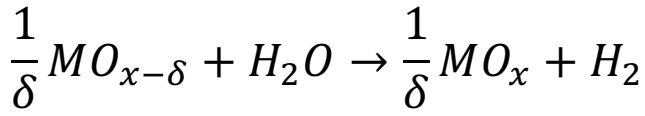
- CeO₂
- Fe(Fe,X)₂O₄
- (A,A')BO₃



Thermal reduction (TR)
High T (~1673 K)
~vacuum ($p_{O_2} \sim 10$ Pa)

Water splitting (WS)
Low T (~873 K)
High H₂ yield ($\frac{p_{H_2O}}{p_{H_2}} = 9$)

State-of-the-art:
Pure fluorite-
CeO₂



Similar cycle for CO₂ splitting

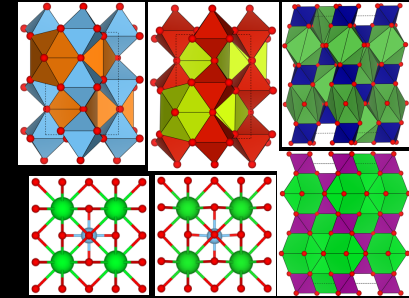
Needs “good” materials!

- Durability (withstand high TR and low WS temperatures)
- Capacity (tolerate high degrees of oxygen off-stoichiometry)
- Stability (no undesired phase transformations)

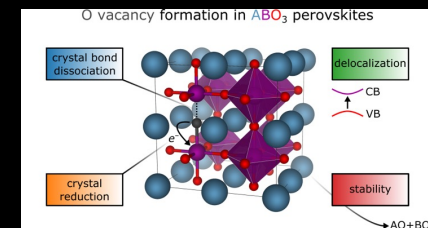
Oxide perovskites: significant current research

Siegel *et al.*, *Ind. Eng. Chem. Res.* **2013**, *52*, 3276
Carillo and Scheffe, *Sol. Energy* **2017**, *156*, 3

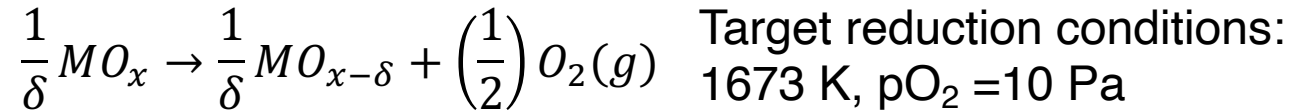
How to use theory to...



- Screen for $\text{H}_2\text{O}/\text{CO}_2$ thermochemical splitters with higher entropy of reduction
- Search through a wider chemical space using machine learning (ML) tools?



Oxygen vacancy formation energy ~enthalpy of reduction



Enthalpy of reduction for the induced off-stoichiometry, δ

$$\Delta H_{reduction} = \frac{H_{MO_{x-\delta}} - H_{MO_x}}{\delta} + \left(\frac{1}{2}\right) H_{O_2}(g)$$

If $\delta \rightarrow 0$, then

$$\Delta H_{reduction} = - \left. \frac{dH_{MO_x}}{dx} \right|_x + \left(\frac{1}{2}\right) H_{O_2}(g) \equiv \Delta H_{formation}^{VaO}$$

Oxygen vacancy formation energy

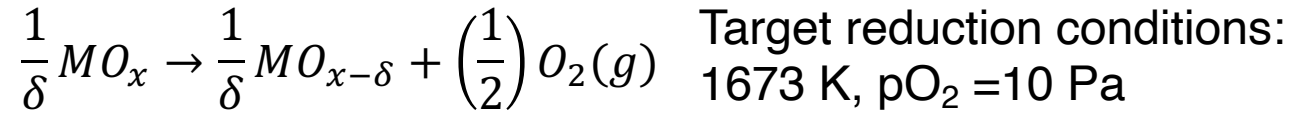
Low $\Delta H_{reduction}$ = large δ

But induced δ needs to be recovered during water-splitting (re-oxidation) \rightarrow optimal $\Delta H_{reduction}$

Thermodynamic models¹ point to optimal $\Delta H_{reduction}$ or $\Delta H_{formation}^{VaO} \sim 3.25$ - 3.5 eV (CeO₂ is ~ 4 eV)

For any theoretical screening, 3-3.75 eV range is ok

Oxygen vacancy formation energy ~enthalpy of reduction



Enthalpy of reduction for the induced off-stoichiometry, δ

$$\Delta H_{reduction} = \frac{H_{MO_{x-\delta}} - H_{MO_x}}{\delta} + \left(\frac{1}{2}\right) H_{O_2(g)}$$

Several studies have theoretically screened, high-throughput calculations and/or machine learning, for novel ABO_3 perovskites, based on the optimal $\Delta H_{formation}^{VaO}$

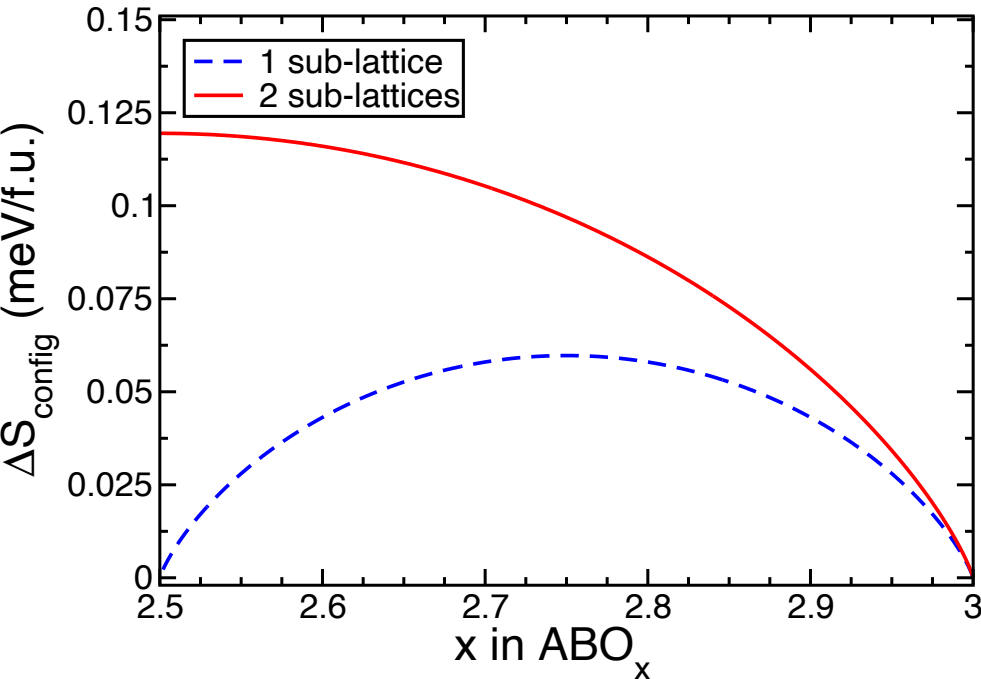
- Candidates either exhibit rare elements (e.g., Eu, Ho) or don't exceed CeO_2 's performance in experiments (stability or kinetic limitations/theory error)
- Need more viable candidates: can we include other metrics?

Thermodynamic models¹ point to optimal $\Delta H_{reduction}$ or $\Delta H_{formation}^{VaO} \sim 3.25$ - 3.5 eV (CeO_2 is ~ 4 eV)
For any theoretical screening, 3-3.75 eV range is ok

Higher entropy of reduction = higher yield

Entropy of reduction for an induced off-stoichiometry, δ , in ABO_3

$$\Delta S_{red} = \underbrace{\frac{S_{ABO_{3-\delta}} - S_{ABO_3}}{\delta}}_{\text{solid}} + \underbrace{\left(\frac{1}{2}\right) S_{O_2}(g)}_{\text{gas}}$$

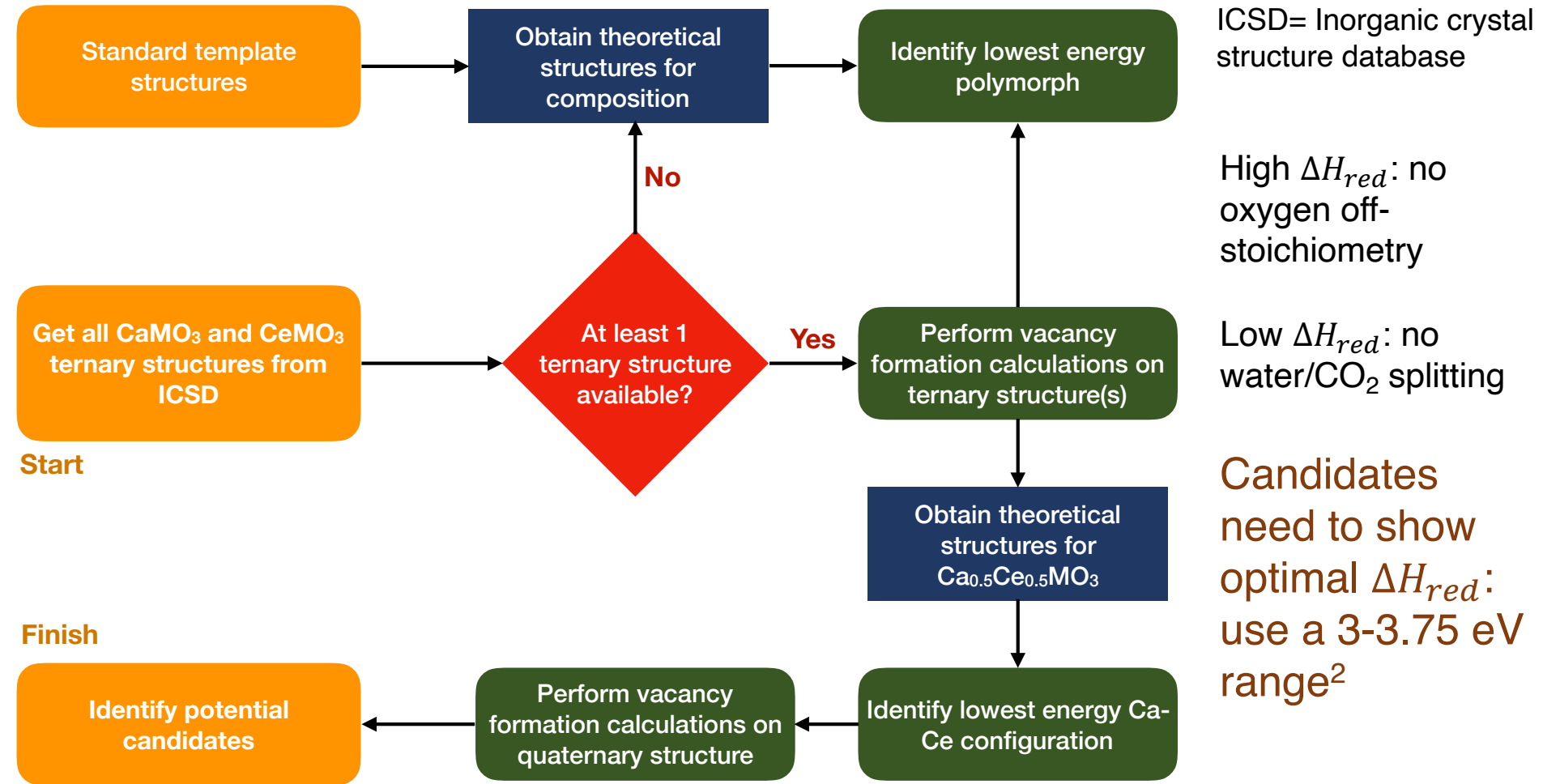


Large contribution to the **solid** portion of ΔS_{red} comes from configurational entropy

Assuming ideal solution of mixing, large increase in ΔS_{red} with A+B cation reduction in ABO_3

For given (optimal) ΔH_{red} , higher ΔS_{red} = higher capacity

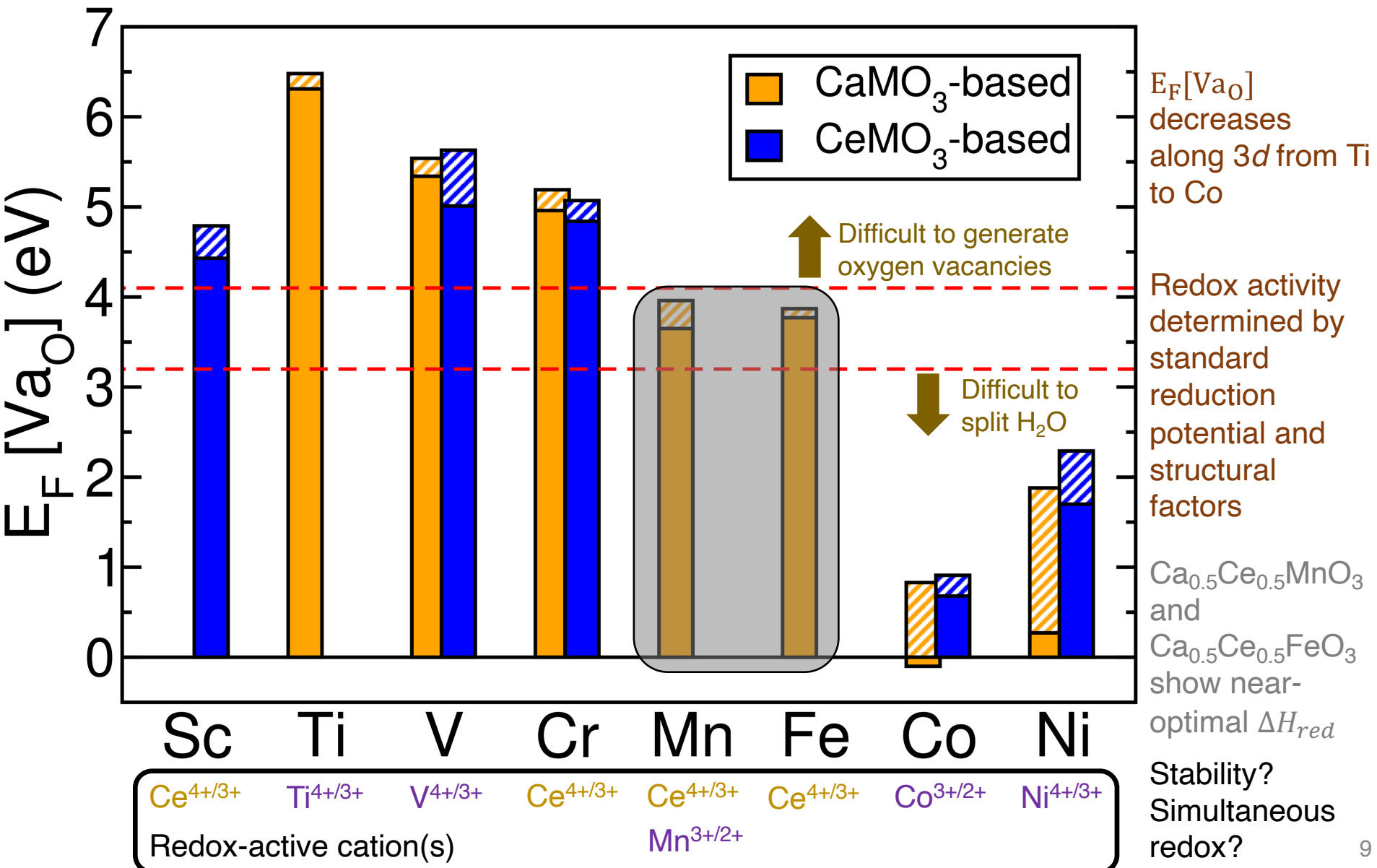
No experimental $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MO}_3$ structures available: use CaMO_3 or CeMO_3



Density functional theory engine: strongly constrained and appropriately normed (SCAN) functional, corrected with optimal Hubbard U correction¹

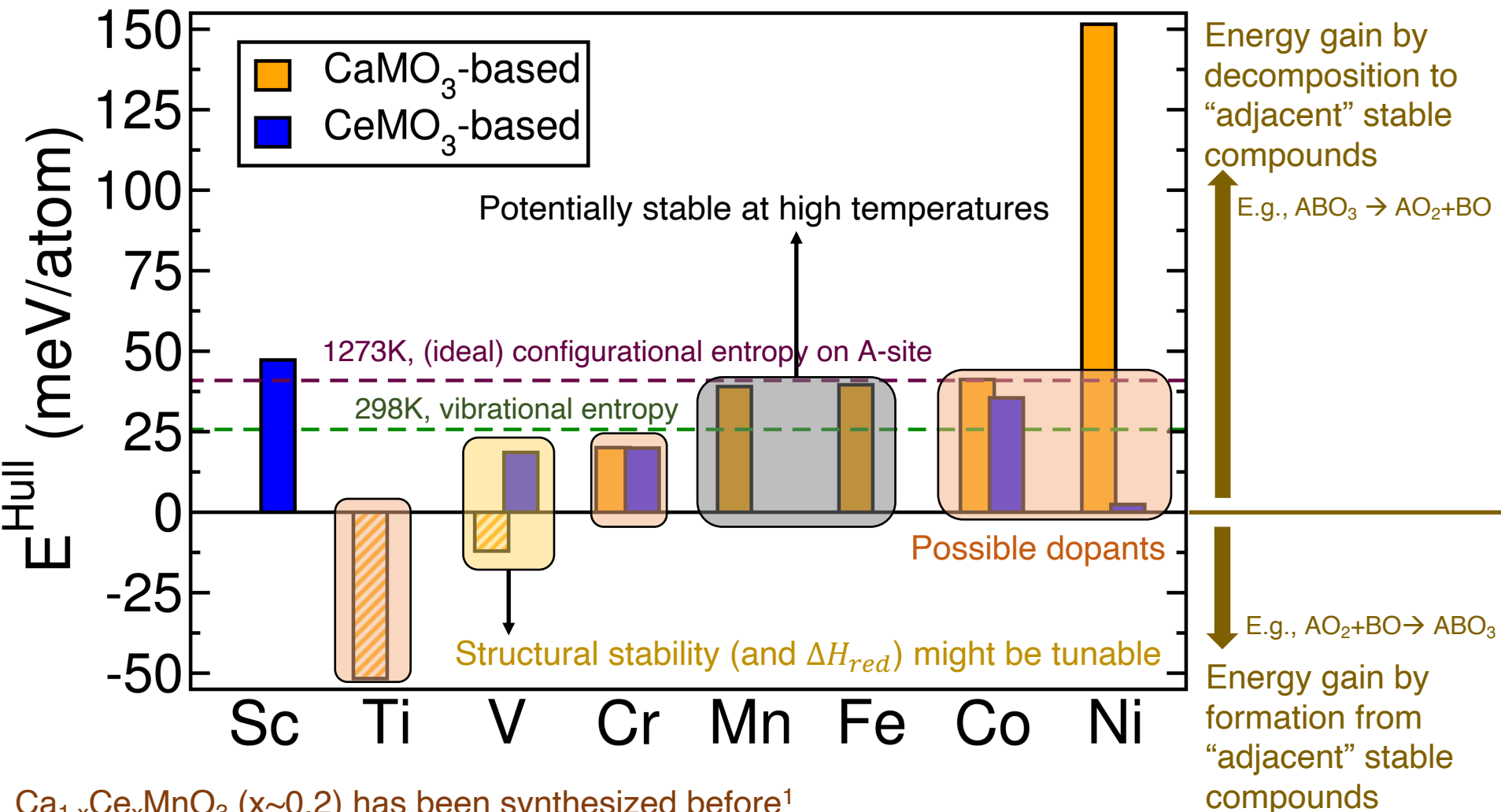
- $\Delta H_{red} \approx E_F[\text{Va}_O]$ (oxygen vacancy formation energy)

Oxygen vacancy formation energy in $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MO}_3$: $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MnO}_3$ and $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{FeO}_3$ are promising



0 K stability of $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MO}_3$

Impact of configurational entropy (of mixing)



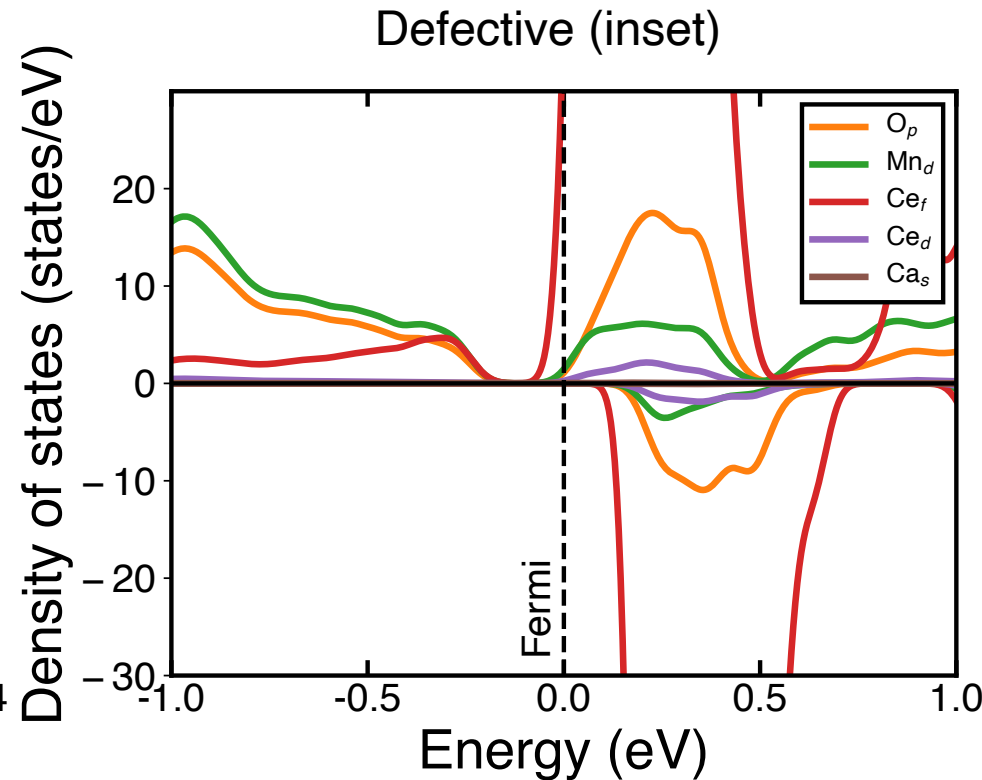
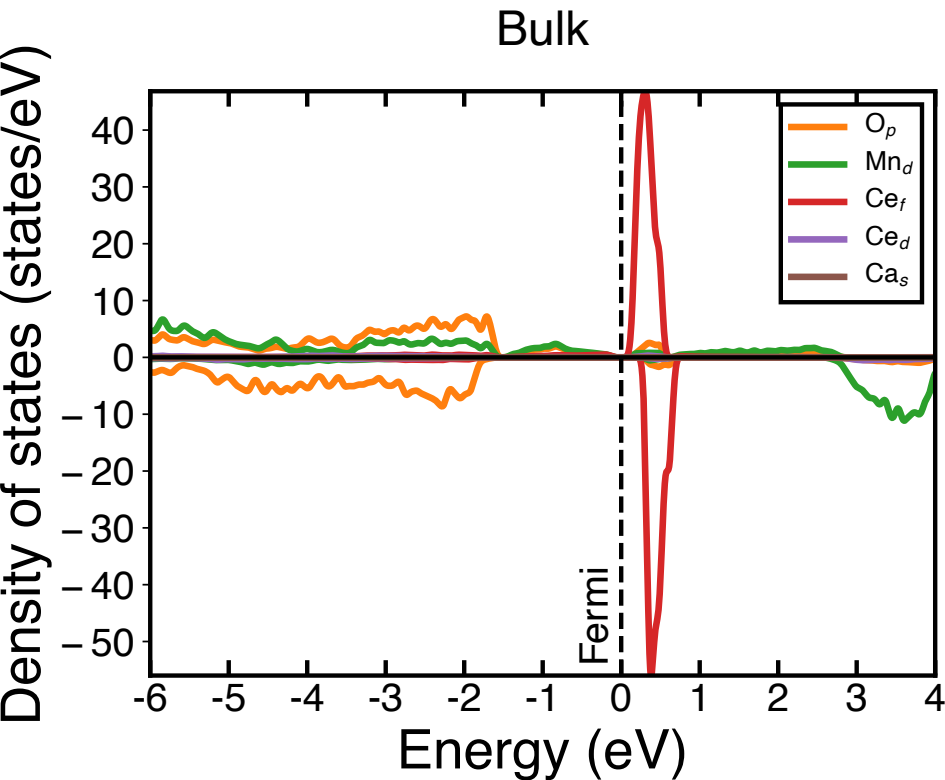
$\text{Ca}_{1-x}\text{Ce}_x\text{MnO}_3$ ($x \sim 0.2$) has been synthesized before¹

$\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MO}_3$ perovskites: can be stabilized at higher temperatures via A-site configurational entropy

1. Zeng *et al.*, *Phys. Rev. B* 2001, 63, 224410

Simultaneous Ce+Mn reduction in $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MnO}_3$

$$E_F[\text{Va}_0] = 3.65\text{-}3.96 \text{ eV} (< \text{CeO}_2, 4\text{-}4.3 \text{ eV}^1)$$

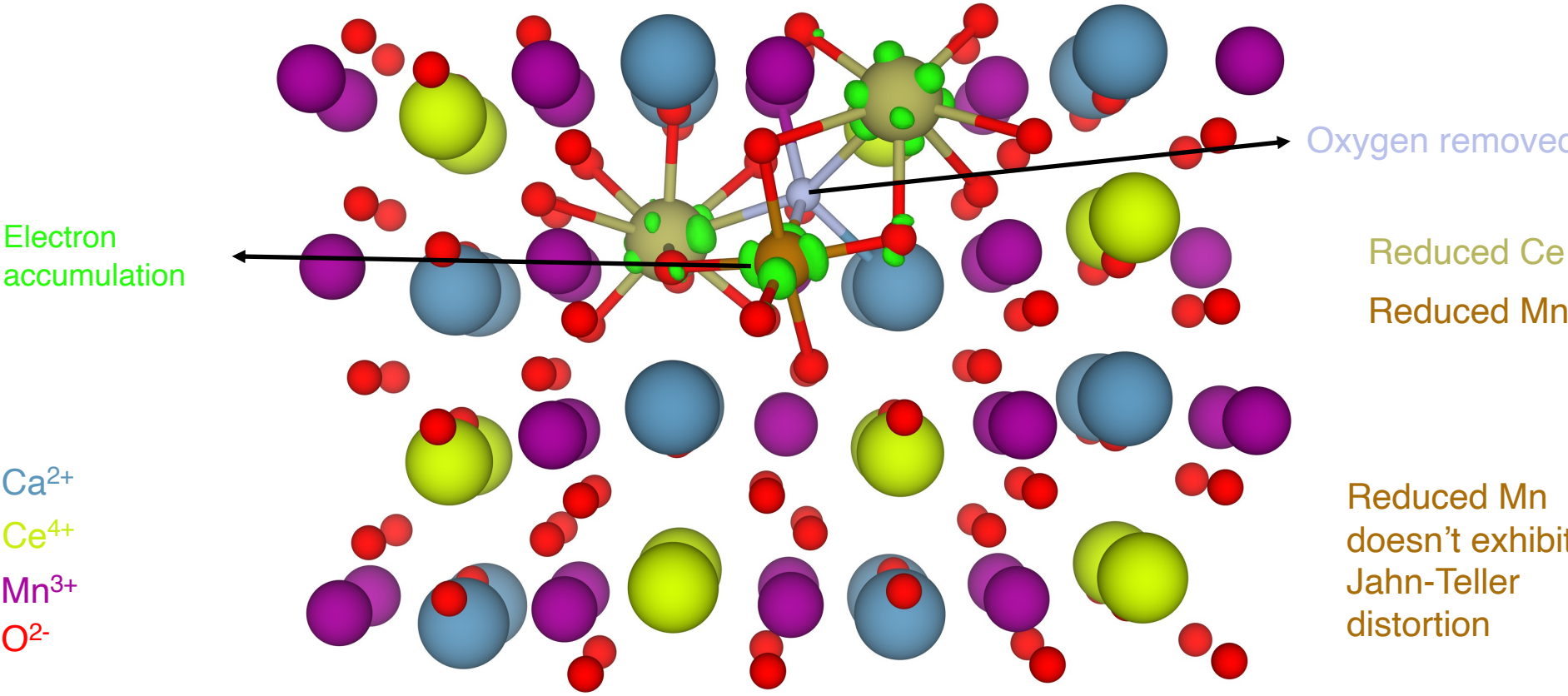


On-site magnetic moments: Ce^{4+} , Mn^{3+} in bulk
 Mn^{3+} displays Jahn-Teller distortion
Ce f beyond Fermi; small amount of Mn d states

On-site magnetic moments: $\text{Ce}^{4+/3+}$, $\text{Mn}^{3+/2+}$
States: Ce f reduction clear, sliver of Mn d

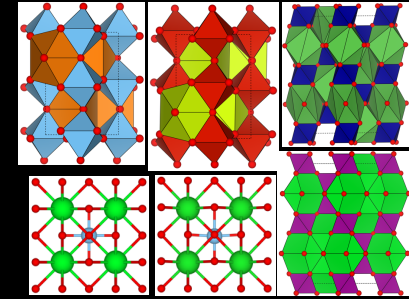
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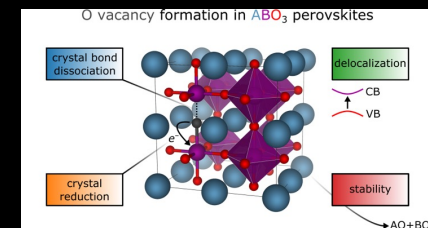


Electron density difference plot: $\text{Ca}_{0.5}\text{Ce}_{0.5}\text{MnO}_3$ with and without oxygen vacancy

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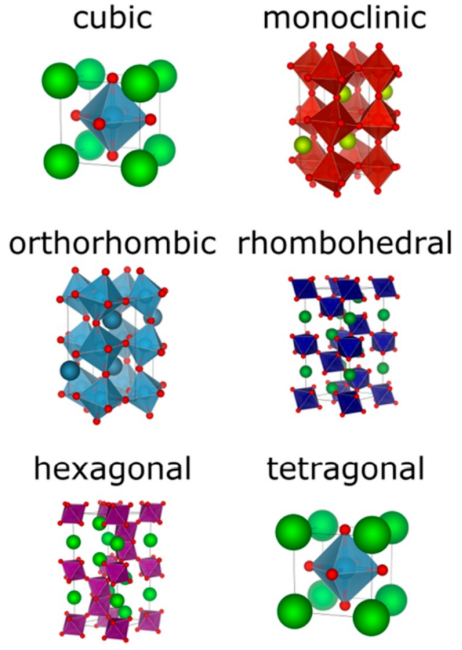
Screening a wider perovskite space for candidates

- Focused computational study on a select class of oxide perovskites: discovery of $(Ca,Ce)MnO_3$ and $(Ca,Ce)FeO_3$
- Possible to identify other candidates?
 - Key property: $E_F[VaO]$
 - Not possible to look for high entropic candidates in a high-throughput fashion
 - Thermodynamic stability (E^{hull}) also important
- Workflow: construct a theoretical (ML) model that can predict $E_F[VaO]$
 - Training the model: generate diverse dataset

H																					
Li	Be																				
Na	Mg																				
K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn										
Rb	Sr	Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd										
Cs	Ba		Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg										
Fr	Ra		Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	Cn										
		La	Ce	Pr	Nd	Pm	Sm	Eu	Gd	Tb	Dy										
		Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf										

■ Alkaline earth A-site 2+ cations
■ Lanthanide A-site 3+/4+ cations
■ 3d transition metal B-site 2+/3+/4+ cations

Redox active → Ce



233 composition-structure combinations, 341 unique $E_F[VaO]$

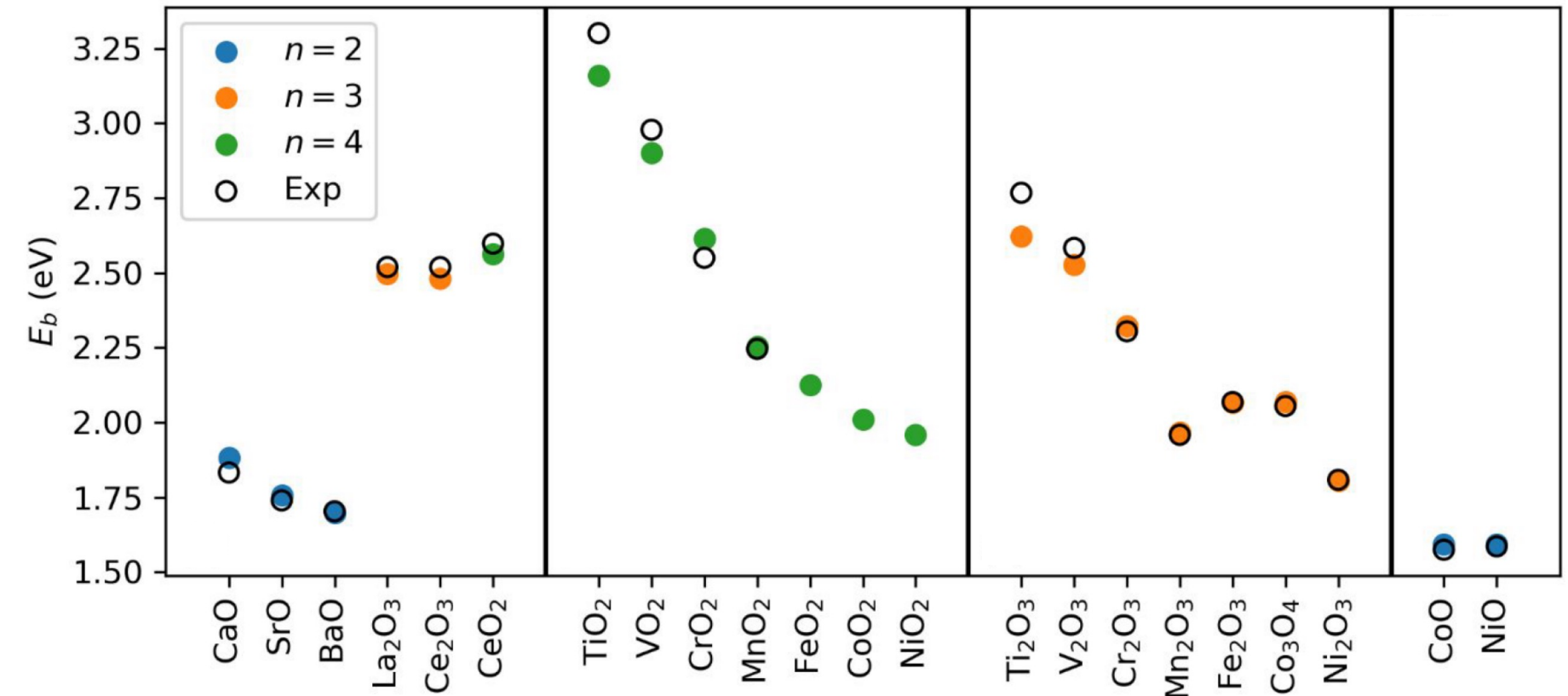
Descriptor for ML model: crystal bond dissociation energy (E_b)

Analogous to molecular bond dissociation energy: energy required to break a metal-oxygen bond in a crystal

$$E_b[O^{2-} - M^{n+}] = \frac{E_c[MO_{n/2}]}{N_b[O^{2-} - M^{n+}]}$$

Cohesive energy of $MO_{n/2}$ oxide

Number of M-O bonds per formula unit in $MO_{n/2}$



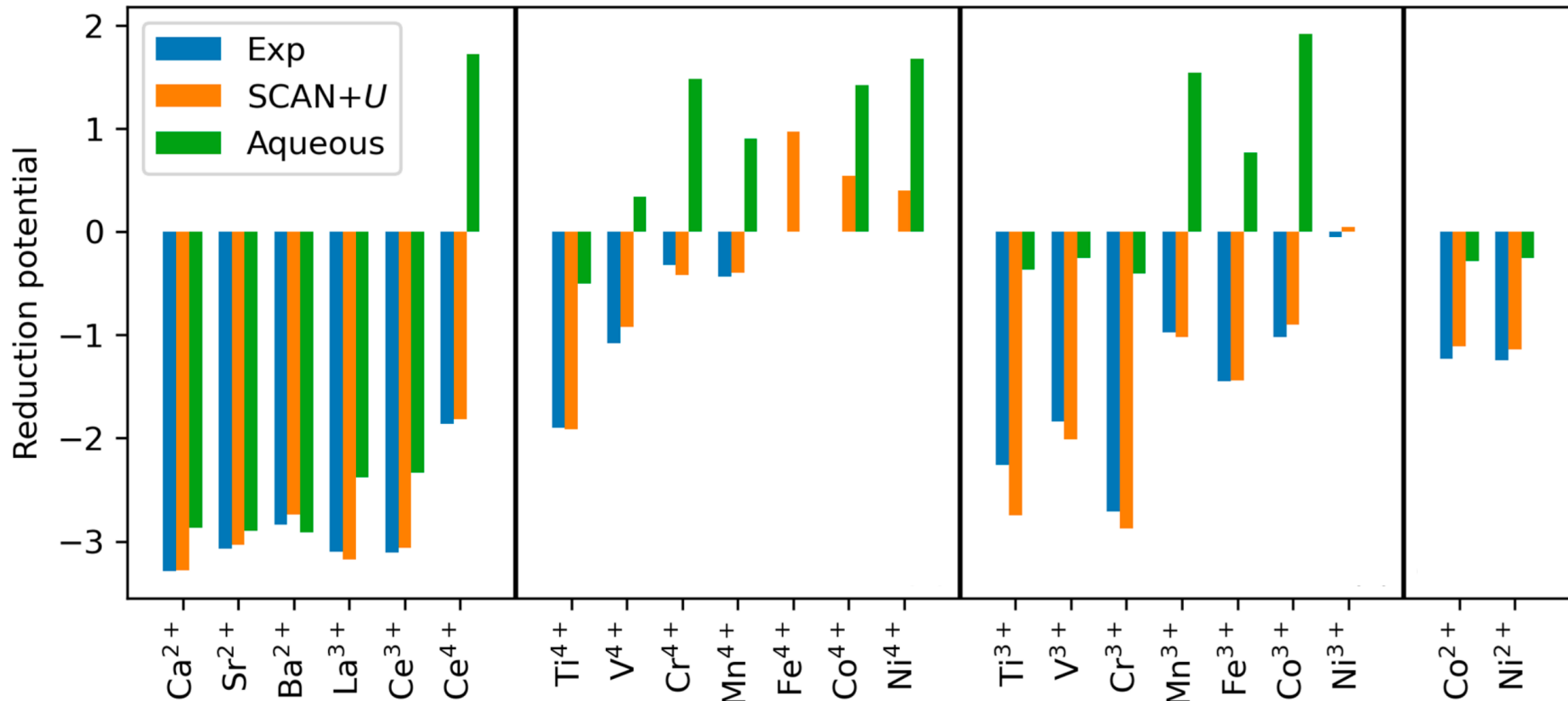
Theoretical E_b in good agreement with experiments

Descriptor for ML model: crystal reduction potential

Oxygen vacancy introduction: reduces (transition-metal) cation, need to account for “ease” of reduction

Standard (aqueous) reduction potentials:

- Are not strictly applicable to solids (no crystal electrostatics included)
- Missing data for few 3d transition metal redox couples



Better systematic trends with crystal reduction than aqueous potential

Descriptor for ML model: thermodynamic stability and band gap

Overall thermodynamic stability of perovskite structure can influence

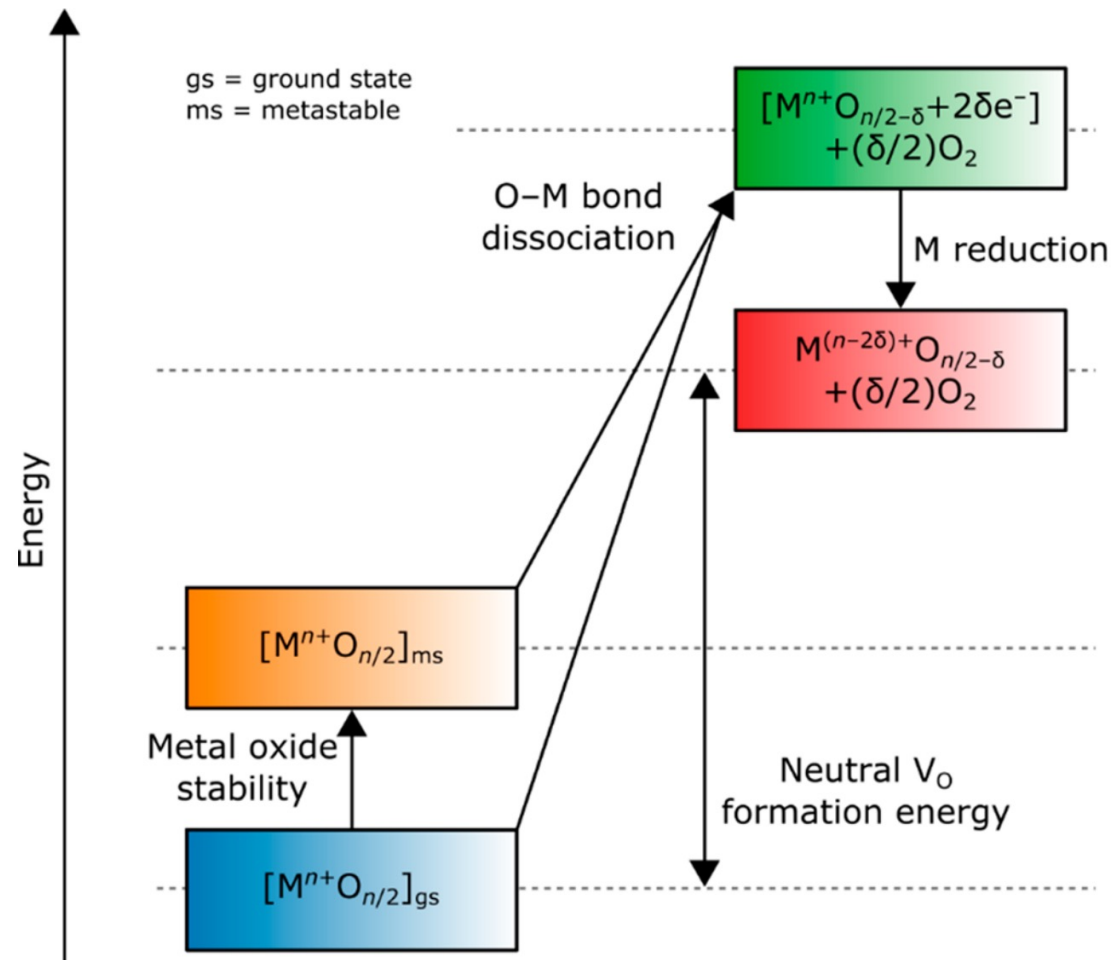
$E_F[Va_O]$

- High stability \rightarrow strong bonding \rightarrow high $E_F[Va_O]$
- High instability \rightarrow weak bonding \rightarrow low $E_F[Va_O]$
- Quantify stability using E^{Hull} at 0 K
- E^{Hull} : includes all possible competing phases

Ease of addition of electrons from oxygen vacancy formation, via delocalization, can alter $E_F[Va_O]$

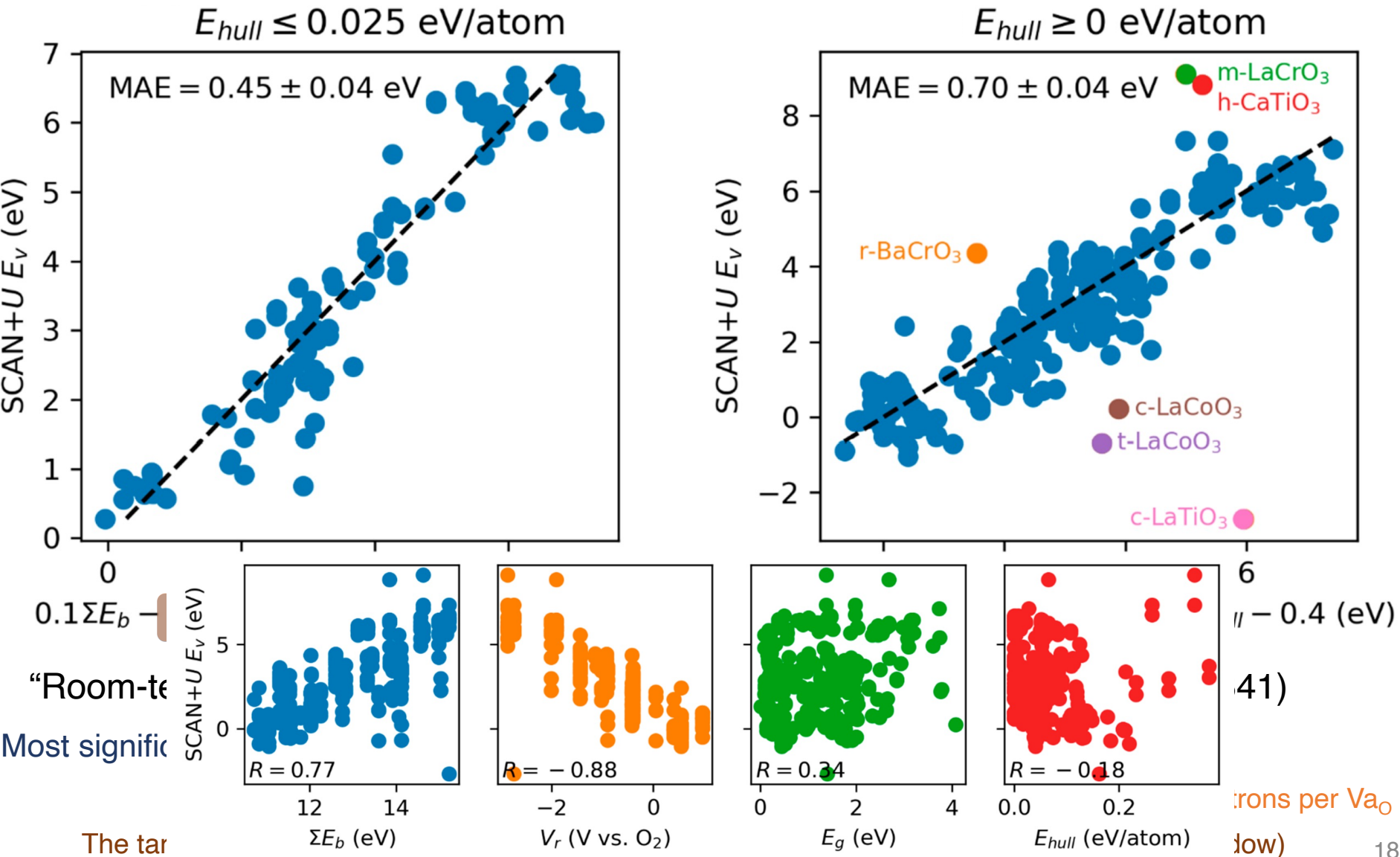
- Large band gap \rightarrow requires E_g (band gap) energy \rightarrow difficult to delocalize
- Small band gap \rightarrow small $E_g \rightarrow$ facile delocalization
- Quantify with E_g at Γ -point
 - Provides upper bound of E_g
 - Computationally not expensive

Descriptors: E_b, V_r, E^{Hull} , and $E_g(\Gamma)$

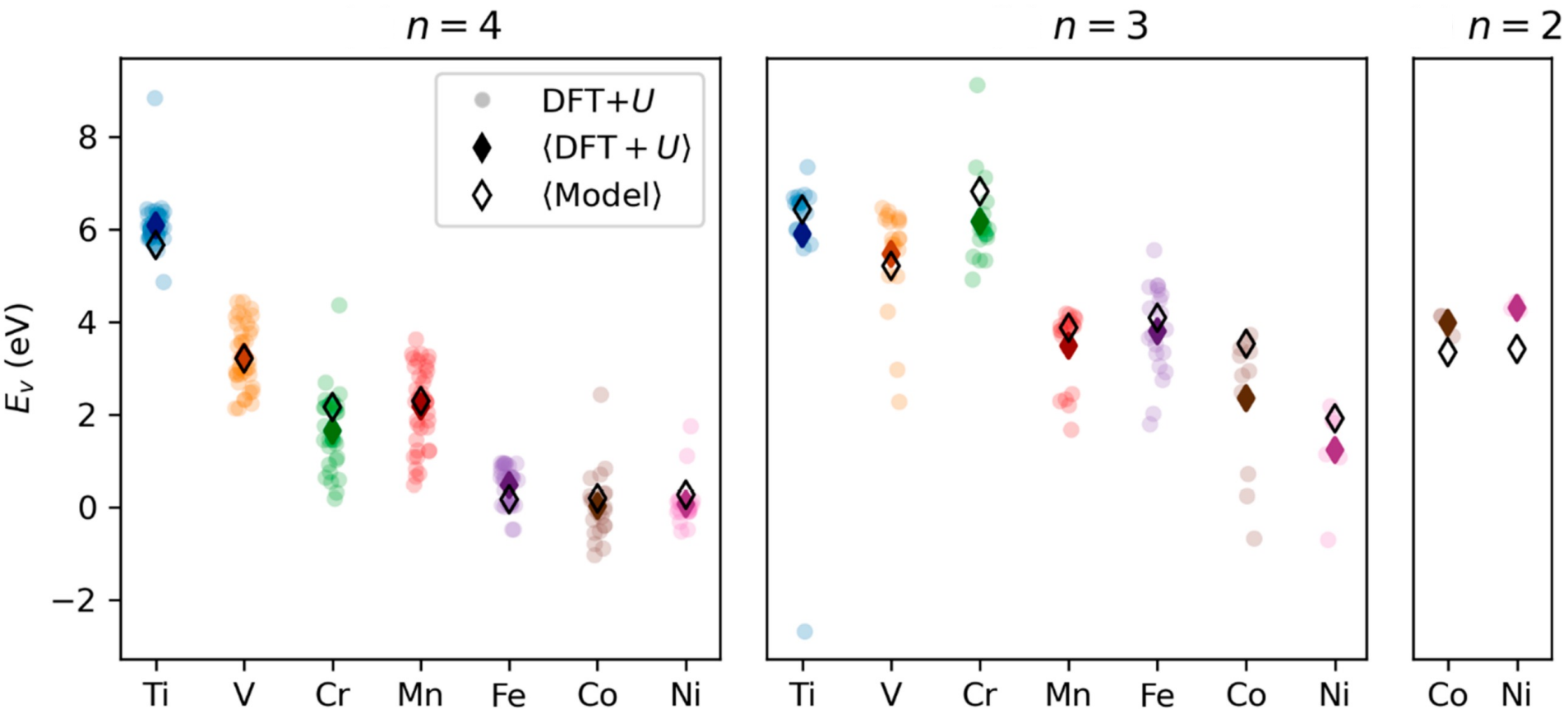


Va_O formation can be thought of as a Born-Haber cycle

Model performance: robust especially for perovskites with low (meta)stability



New candidates



Trends for low/intermediate/high E_v : identify candidates

Possible B-site cations for solar thermochemical (3-3.75 eV): V^{4+} , Mn^{3+} , Fe^{3+} , $\text{Co}^{3+/2+}$, and Ni^{2+}

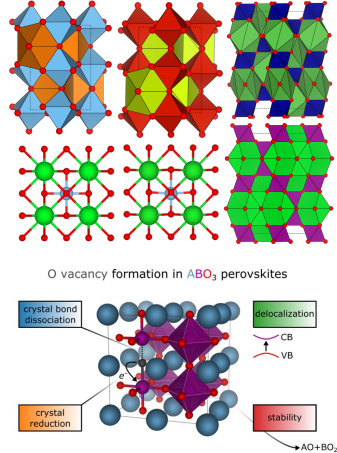
Candidates from outside the training set:

- $R\bar{3}c$ - BiFeO_3 (model $E_v = 3.99$ eV)
- $P4mm$ - BiCoO_3 (model $E_v = 3.80$ eV)

Experimental validation of new candidates in progress 19

Conclusions and Acknowledgments

- Need better materials for STC H₂O/CO₂ splitting
 - Oxide perovskites: tunable ΔH_{red} , low ΔS_{red}
 - Identify simultaneously redox-active perovskites with optimal ΔH_{red} (3-3.75 eV) to improve ΔS_{red}
 - Size + charge-neutrality + redox-activity constraints = Ca_{0.5}Ce_{0.5}MO₃ (M = Sc, Ti, ..., Ni)
- Ca_{0.5}Ce_{0.5}MnO₃ and Ca_{0.5}Ce_{0.5}FeO₃ exhibit near-optimal ΔH_{red} ($\approx E_F[Va_O]$)
 - Most quaternaries are not stable at 0 K, but A-site configurational entropy can help
 - Ca_{0.5}Ce_{0.5}MnO₃: simultaneous redox-active candidate
- Built a machine learned model on a diverse set of perovskite structures to predict Va_O formation energy
 - Two candidates from outside the training set: BiFeO₃ and BiCoO₃

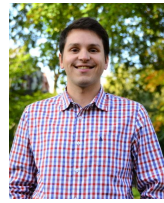


Theoretical screening:

“Exploring Ca-Ce-M-O (M=3d transition metal) oxide perovskites for solar thermochemical applications”, G.S. Gautam, E.B. Stechel and E.A. Carter, **Chem. Mater.** **2020**, *32*, 9964-9982

Machine learning model:

“Factors governing oxygen vacancy formation in oxide perovskites”, R.B. Wexler, G.S. Gautam, E.B. Stechel, and E.A. Carter, **J. Am. Chem. Soc.** **2021**, *143*, 13212-13227



Tiger (Princeton)



NREL (Eagle)



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